SOLAR DIVISION BULLETIN.

Neal J.Hoines, Editor.

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560 Broadway. Paterson 4. Now Jorsey.

The average length of a sunspot cysle is 11.5 years. The period from minimum to maximum is four to five years and from maximum to minimum seven to nine years.

In the present eyelo we are in the minimum to maximum period. This period has been unusually active although we are now only two and one-half years past minimum. Because of this unusual activity

Research prodicts an early maximum.

With the approach of maximum we would like the S.D. observers to take special notice of the frequency of Peres, Veiled Spets, Feint Markings, Ponumbral Wieps, and, Faculacic Tracory. Kindly enter data of same in the Romarks Column or on the back of the Monthly Report form.

PORES.

Porcs are small, short-lived, spots, without penumberal areas obsorved mostly in the Central Zone of the sun. Some are visible for about a day but mostly they last only a few hours. VEILED SPOTS!

Voiled spots are similar to peres but seem to be below a hazy cloud instead of on the surface as are the pores. Like the pores

they are observed mostly, in the Central Zone. FAINT MARKINGS.

Faint Markings are observed within the confines of both sunspot belts. They appear between, in, above and below groups and are seen entirely alone, often in the Central Zone. They are tiny ponumbral areas or a small group of tiny spots close together. Occasionally they develope into an active grpup. When once observed, that area should be closely watched for now group activity. PENUMBRAL WISPS.

Penumbral Whisps are mostly seen near "F" type groups. In appearance they resemble small penumbral areas which have broken away from the main body of the "F" group. Semetimes they appear like whirling patches of penumbra. FACULARIC TRACERY.

This is a coined word by your chairman. It is so used to dosignate a trace of faculae. We are familiar with faculae as seen near the east and west limbs of the sun, around, and adjacent to sunspet groups. Faculacic Tracery is seen only in the Central Zone, has the same appearance as faculae but very faint and diffused. When observed it is seen nestly alone.

All of the above items are observed visually with a Herschelian Eyepicco. It is doubtful that the projection nothed of observation will rovoal any of the above items except possibly the percs. Observational proceedure used here is as follows. Adjust R.A. Clamp for free movement of the telescope tube by hand. When seen or suspected the tube is moved, slowly, to the right and left as this has a tendency to bring out irregularities on the solar surface.

In your Instructions Leaflet, page one, fifth paragraph, the last sentence, we find. "Entries should be made with a hard pencil as the sharp point allows more detail, especially where an entry embodies multiple date ". With more groups and spots now in evidence this instruction should be closely followed. Some of our observers have an abundance of room on the Monthly Report form while others using dull pencils or pens find it difficult to include the break-down items.

Incidently, it is not necessary to show the totals on the last line of the report form for break-down data in headings "g" and "i" but it is required that you show totals on the last line for headings b,f,g,h,i,j,k,l,m,n,o,p. In addition the sum of j and k should equal the sum of f and likewise the sum of l and m should equal the sum of

g. Study paragraph two on page seven of the loaflet.

If during sunspot observations, color, other than that of the usual color of penumbral or umbral pertions of groups is seen make a separate record of this, and send same to Pr.J.C.Bertlett of 300 N. Eutem Street, Beltimore 1, Ma., aththe end of each month. We have asked Dr.Bartlett to make a special study of this in order to

determine the possible causes of this phonomena.

We repeat an earnest request, that you make every effort to mail your completed reports so that they reach this efffice not later than the tenth of each steeceding menth. Most reports are received on i time even those from foreign countries, but a few are late. A new International set—up has been arranged and the Bureau of Standards in Washington where our reports are new received also has a time limit. Recognise the fact that your data is being used in this new International set—up. The importance of your work and co-operation is self-evident.

From time to time you will receive, through the Bulletin, information concerning the availability of Reprints on the subject in which we are interested. This month we recomend that you write to Dr.Ralph E.DeLury, Dominion Observatory, Ottawa, Canada, for their Dom. Obs. Reprint, No 32, "Sunspot Influences ". With your request include about 15 cents for postage and mention the Solar Division as your source of information. (Do not send stamps).

Observers with Reflecting Telescopes are urged to participate in the color study of sunspets. We are in need of some visual observations with this instrument. Here is the possibility of a good check on effects of secondary spectrum in achromatics. Arrange your

data and sond same to Dr.Bartlett.

We are happy to add to our nombership list Mr. Joseph Swaelen of Bolsuim, J. Hillobrand of Detroit and Floyd Arnold Arnold of Brooklyn.

osting photos of his new mounting for projection observing with a

Refracting telescope . Cant stop on folks.

The fall meeting of the AAVSO will be held as usual in October at Harvard College Observatory, Cambridge, Mass., make every sacrifice to attend. They come from miles and niles and leave with smiles. Each meeting brings a better representation of the Solar Diviison. Reverend Kearens, our S/D/ COMMITTEE member and our "top man" in Solar Photography writes that he has a complete series of pictures of 46117 the great sunspet group of January and February. We received some prints of the March return of theis group from Rev. Kearens a few days ago. Reverend Kearens will exhibit these pictures at the fall meeting. Meet the AAVSO folks this fall, a wenderful experience.