UPDATING THE PERIOD OF V COMAE BERENICES

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ABSTRACT: In AAVSO Abstracts, 19 October 1968, Jean Jackson pointed out that the period of V Com is changing. Since then additional observations have been acquired from Harvard, Nantucket, and Yale plates, and more published observations have been compiled. All this material was turned over to Janet Johnston for a re-evaluation of the period. Her results, based on some 450 miscellaneous observations, mainly between JD 18443 and 42188, indicate that the observed phases, Ø, are satisfactorily represented by the relation,

$$\emptyset$$
 + n = (JD - JD_O)P_O⁻¹ + kn², where JD_O = 19000 P_O^{-1} = 2.131567, or P_O = 0^d.469 1384 and k = 1.7 x 10⁻¹⁰.

It is noted that the basic period, P_0 , agrees with the period as given in the General Catalogue of Variable Stars. The correction term implies a progressive shortening of the period (dP/dn = $2kP_0$) amounting to 1.6×10^{-10} day/cycle. This is reasonably consistent with the difference between the basic periods given in the GCVS and by Miss Jackman, which refer to different basic epochs, JD_0 .

D.H.

THREE VARIABLE STARS IN CYGNUS

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ABSTRACT: Lucia Dexter, a summer undergratuate assistant at the Maria Mitchell Observatory, reported primarily on a new RR Lyrae type star at 19^h36^m39^s+38^o44'.7 (1900), discovered in 1973 by Bonnie Buratti. Step estimates of its brightness were made by Miss Dexter on some 900 Nantucket plates taken between JD 24695 and 42314. Over this long time interval she found the period to be constant at 0^d.581 147.

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A new eclipsing variable found by Miss Dexter at 19^h24^m34^s +41^o20'.9 has an approximate period of 3^d.72, but she did not have time available to complete the estimates. More definitive results on both of these stars will be published later when further work on them has been completed.

The third star on which Miss Dexter reported was HQ Lyr, which she had rediscovered, and for which she found a provisional period of 290 days (see IBVS No. 950, Dec. 1974).

D.H.